

The Evolution of Galaxies

II-Basic Building Blocks

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and D. Schuster



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**Marc Sauvage, Grazyna Stasinska, D.
Schaerer**



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The Evolution of Galaxies Marc Sauvage, Grazyna Stasinska, D. Schaerer, 2013-06-29 Galaxies have a history This has become clear from recent sky surveys which have shown that distant galaxies formed early in the life of the Universe differ from the nearby ones New observational windows at ultraviolet infrared and millimetric wavelengths provided by ROSAT IRAM IUE IRAS ISO have revealed that galaxies contain a wealth of components very hot gas atomic hydrogen molecules dust dark matter A significant advance is expected due to new instruments VLT FIRST XMM which will allow one to explore the most distant Universe Three Euroconferences have been planned to punctuate this new epoch in galactic research bringing together specialists in various fields of Astronomy

The Evolution of Galaxies G. Hensler, G. Stasinska, S. Harfst, P. Kroupa, C. Theis, 2013-11-11 Galaxies have a history This has become clear from recent sky surveys which have shown that distant galaxies formed early in the life of the Universe differ from the nearby ones New observational windows at ultraviolet infrared and millimetric wavelengths provided by ROSAT IRAM IUE IRAS ISO have revealed that galaxies contain a wealth of components very hot gas atomic hydrogen molecules dust dark matter A significant advance is expected from the results of new instruments VLT FIRST XMM which will allow one to explore the most distant Universe Three Euroconferences were planned to punctuate this new epoch in galactic research bringing together specialists in various fields of Astronomy This book contains the proceedings of the third conference and presents the actual state of the art of modelling galaxy evolution

The Evolution of Galaxies, 2002 *Galaxies: the Third Dimension* Margarita Rosado, Luc Binette, Lorena Arias, 2002 Transactions of the International Astronomical Union International Astronomical Union, 2003 Kapteyn Astronomical Institute Annual Report, 2001 *Annual Report* Kapteyn Instituut, 2002 Information Bulletin International Astronomical Union, 2000 *NIC Symposium 2001* John von Neumann-Institut für Computing. Symposium, 2002

A Massive Star Odyssey International Astronomical Union. Symposium, 2003 *Reviews in Modern Astronomy* Reinhard E. Schielicke, 2002-12-10 This 15th volume in the annual series on recent developments and scientific progress in astronomy and astrophysics contains fourteen invited reviews presented during the Joint European and National Astronomical Meeting JENAM 2001 held in Munich Germany Readers also learn about the lecture on macro and microscopic views of nearby galaxies given by Keiichi Kodaira Japan who was awarded the Karl Schwarzschild medal 2001 Further contributions on the topic provide the latest results on the search for extra solar planets formation of stars and galaxies physics of active galactic nuclei as well as new telescopes and sensor technologies for various wavelengths

Mitteilungen der Astronomischen Gesellschaft Astronomische Gesellschaft (Germany), 2002 **Tatigkeitsbericht des Astronomischen Rechen-Instituts in Heidelberg** Astronomisches Rechen-Institut, Heidelberg, 2001 **Evolution of Galaxies** M. Sauvage, G. Stasinska, D. Schaerer, 2002 Reprint, 1958 **International Aerospace Abstracts**, 1995 **Galaxies**, 1995 Secular Evolution in Galaxies Justus Neumann, 2019 Galaxies are gravitationally bound systems of stars gas dust

and probably dark matter. They are the building blocks of the Universe. The morphology of galaxies is diverse; some galaxies have structures such as spirals, bulges, bars, rings, lenses, or inner disks, among others. The main processes that characterise galaxy evolution can be separated into fast, violent events that dominated evolution at earlier times and slower processes which constitute a phase called secular evolution that became dominant at later times. Internal processes of secular evolution include the gradual rearrangement of matter and angular momentum, the build up and dissolution of substructures, or the feeding of supermassive black holes and their feedback. Galaxy bulges, bright central components in disc galaxies, on one hand are relics of galaxy formation and evolution. For instance, the presence of a classical bulge suggests a relatively violent history. In contrast, the presence of a disc-like bulge instead indicates the occurrence of secular evolution.

On the Evolution of Massive Galaxies Kristen Leah Shapiro, 2010. Galaxies as the fundamental building blocks of the Universe are the critical link between the overall evolution of the Universe and the assembly of small-scale structures such as stars and planets within it. Unfortunately, the formation and evolution of galaxies remains poorly understood due to the incredible complexity of the physics that governs these processes. To study and constrain these processes, a particularly useful galaxy population are bulge-dominated galaxies such as elliptical and large spiral galaxies, which together are the most massive and most evolved components of the local Universe. In the present day, these galaxies are dominated by old stars; however, their histories likely include an epoch of powerful star formation and rapid growth of their supermassive black holes. Progress in understanding the evolution of massive galaxies can therefore proceed on two fronts: 1) observations of their formation in situ in the early Universe, and 2) detailed studies of the fossil relics of this process in the local Universe, with the ultimate goal being to link progenitors and descendants. A key epoch for such investigations is 10 billion years ago, the most active period in the Universe's history at which time the vast majority of stellar material in galaxies was assembled. Recent comparisons of the observed properties of galaxy populations across cosmic time have shown that the dominant star-forming galaxy population at these early times were the probable ancestors of present-day massive bulge-dominated spiral and elliptical galaxies. The obvious direction for current and future research is therefore to probe the detailed evolution with time of the properties and substructures that define this local galaxy population. This goal has guided my dissertation research as described in the following pages. Using photometric, spectroscopic, and integral field observations at optical through mid-infrared wavelengths, I have studied both star-forming galaxies in the early Universe and their present-day descendants. Specifically, this thesis explores the dynamical star-forming and black hole properties of galaxies 10 billion years ago and shows that these young galaxies must be assembled via a rapid but steady influx of gas from the surrounding cosmic structure. The resulting large quantity of gas in these galaxies causes super-large star-forming gas clouds to form, and the dynamical interactions of these clouds control the evolution of the galaxies' supermassive black holes and internal substructures, producing the bulges and globular cluster populations observed in the present day. Studies of the resulting local massive galaxy population also

presented herein confirm that such successive minor dynamical disturbances were important to the assembly of these bulge dominated galaxies and their supermassive black holes In the pages of this thesis an exciting link is emerging in which many observed properties of local galaxies can be explained by the dramatic internal processes occurring in galaxies 10 billion years ago during the era of the most rapid galaxy assembly *Resolving the Building Blocks of Galaxies in Space and Time*
,2017

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